

M-theory Potential from the G_2 Hitchin Functional in Superspace

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Abstract

We embed the component fields of eleven-dimensional supergravity into a superspace of the form $\mathbf{X} \times Y$ where \mathbf{X} is the standard 4D, $N = 1$ superspace and Y is a smooth 7-manifold. The eleven-dimensional 3-form gives rise to a tensor hierarchy of superfields gauged by the diffeomorphisms of Y . It contains a natural candidate for a G_2 structure on Y , and being a complex of superforms, defines a superspace Chern-Simons invariant. Adding to this a natural generalization of the Riemannian volume on $\mathbf{X} \times Y$ and freezing the (superspin- $\frac{3}{2}$ and 1) supergravity fields on \mathbf{X} , we obtain an approximation to the eleven-dimensional supergravity action that suffices to compute the scalar potential. In this approximation the action is the sum of the superspace Chern-Simons term and a superspace generalization of the Hitchin functional for Y as a G_2 -structure manifold. Integrating out auxiliary fields, we obtain the conditions for unbroken supersymmetry and the scalar potential. The latter reproduces the Einstein-Hilbert term on Y in a form due to Bryant.

Contents

1	Introduction	1
2	Superfields and Components	4
3	Action	7
3.1	Chern-Simons Action	7
3.2	Kähler Action	10
4	Scalar Potential	12
5	Discussion	13
A	G_2 Toolbox	16

1 Introduction

In the zoo of supergravity theories, eleven-dimensional supergravity is unique in that it has the largest possible (manifest) spacetime symmetry group. Despite being, in this sense, the most fundamental of supergravity theories, it has various quite mysterious properties. For example, in contrast to its ten-dimensional relatives, there is no theory of critical superstrings that has it as a low-energy limit. To find a home even somewhat analogous, one must go to M-theory (which is even more mysterious) and take a massless limit of that. Another presumably related property is the emergence of an exceptional symmetry of its (gauged) compactifications on tori.

For applications to the study of physics in lower dimensions, this theory may be compactified on eleven-dimensional manifolds of the form $X \times Y$ and expanded in Kaluza-Klein modes by integrating over Y . This results in an effective theory on X in which the contribution of the internal part is organized in a tower of ever-more-massive fields.

An alternative to this approach is to split the eleven-dimensional spacetime as $X \times Y$ and to reorganize the fields into representations of the reduced structure group but without averaging over the “internal” space. Such backgrounds were precisely the subject of reference [1], wherein this is referred to as “keeping locality in Y ”. There, the action for the bosonic part of eleven-dimensional supergravity was decomposed on $X \times Y$ explicitly. Of course it is always possible to keep the full diffeomorphism invariance of the eleven-dimensional theory recast in terms of covariant, interacting X and Y parts. What

is somewhat surprising, however, is that this can be organized in a very manageable form [1]. We would like to construct a superspace action that reproduces the bosonic eleven-dimensional supergravity action in this form.

As this is presumably impossible (in the naïve sense) for more than 8 real supercharges, we settle for a superspace description with at most $N = (1, 0)$ supersymmetry in 6D, $N = 1$ in 5D, or $N = 2$ in 4D. These maximal off-shell cases require an infinite number of auxiliary fields [2] and non-chiral matter. This complicates the use of such a superspace description both technically and phenomenologically. Instead, we propose to embed the components of eleven-dimensional supergravity into 4D, $N = 1$ superfields with arbitrary Y -dependence. This gives a description of eleven-dimensional supergravity on $\mathbf{X} \times Y$ with \mathbf{X} a curved superspace modeled on $\mathbf{R}^{4|4}$ and Y a Riemannian 7-manifold. Projecting such a theory to component fields results in a component supergravity description on the bosonic submanifold $X \times Y$.

Although the resulting physics is eleven-dimensionally super-diffeomorphism invariant, only the 4D, $N = 1$ part of the local super-Poincaré symmetry would be manifest (together with the 7D (bosonic) Riemannian part). Note that precisely this amount of local super-Poincaré invariance is what one would retain were one to compactify on a manifold Y admitting a Riemannian metric of G_2 holonomy. Although we do not insist on such a background in this work, it will be useful to adopt the language of 4D, $N = 1$ compactifications in which we refer to X or \mathbf{X} as “spacetime” and Y as the “internal space”.

There are various partial realizations of this superspace supergravity program less ambitious than the construction of the full 11D theory in arbitrary $\mathbf{X} \times Y$. For example, one could attempt to build the linearized action by working out the linearized superdiffeomorphisms and building an invariant action order-by-order following a superspace Noether procedure.¹

Alternatively, one may attempt to define the theory in a gravitino superfield $\Psi(x, y, \theta, \bar{\theta})$ expansion keeping only the 4D, $N = 1$ supergravity fields and the superfields holding the components of the 3-form but all non-linearly. In such an approach, we expect the action to take the form

$$S = S_{CS} + S_K + \mathcal{O}(\Psi) \tag{1.1}$$

¹In the analogous problem for 5D, $N = 1$ supergravity in 4D, $N = 1$ superspace, this was the approach taken in [3, 4]. This setup is related to the eleven-dimensional version considered here by taking $Y = \mathbf{R} \times Y'$ in the massless limit, where Y' is a compact Calabi-Yau 3-fold. This superspace was used in [5] to compute supergravity loop corrections to supersymmetry breaking in (a phenomenological analog of) heterotic M-theory on Y' .

to lowest order. Here, the Chern-Simons action S_{CS} is taken to be the invariant of the non-abelian tensor hierarchy constructed in [6, 7]. (This hierarchy encodes the components of the dimensionally-decomposed eleven-dimensional 3-form.) We will refer to the remaining terms S_K as the “Kähler action”. We propose to take it to be a natural generalization of the super-volume on $\mathbf{X} \times Y$ constructed from the remaining supergravity and tensor hierarchy fields (cf. eq. 3.8).

Additionally, one may consider freezing the 4D, $N = 1$ supergravity multiplet around a flat $\mathbf{R}^{4|4}$ background and letting only the tensor hierarchy fields fluctuate. (We take the spacetime part to be flat for simplicity, but a curved rigid background may be considered instead.) In this approximation, the Kähler term reduces to the superfield analog of the Riemannian volume on Y . In our approach, the metric scalars are the imaginary part of chiral scalar fields in the tensor hierarchy that are 3-forms on Y . (The real part holds the 3-form scalars.) This defines a G_2 structure on Y . The result of this is that the Kähler action is, essentially, a superspace lift of the Hitchin functional [8, 9].

In this work, we test this proposal by computing the scalar potential of this action. As we are freezing the spacetime supergravity part, this background will be of the form $\mathbf{R}^{4|4} \times Y$ with Y a G_2 structure manifold (not necessarily compact). The remaining fields are those of the non-abelian tensor hierarchy. In particular, we reproduce the scalar potential of eleven-dimensional supergravity from the Chern-Simons action and the Hitchin functional. This potential consists of the Ricci scalar on Y in a form due to Bryant [10] and an analogous expression for the square of the Maxwell-like tensor for the 3-form scalars.

Before concluding, let us pause to compare the proposed set-up to the analogous construction for ten-dimensional, $N = 1$ super-Yang-Mills worked out by Marcus, Sagnotti, and Siegel [11]. In that case, the superspace is of the form $\mathbf{R}^{4|4} \times Y'$ where Y' has a fixed $SU(3)$ structure $(\bar{\partial}, \Omega)$. The components of the ten-dimensional gauge field are embedded in a real superfield (1-form components along X) and three chiral superfields ϕ_i (components along Y') transforming in the $\mathbf{3}$ of $SU(3)$. All four superfields are valued in the adjoint representation of the gauge group G . The superpotential of the theory is the superspace generalization of holomorphic Chern-Simons functional

$$W_{MSS} = \frac{1}{2} \int d^4x \int d^2\theta \int_Y \bar{\Omega} \wedge \text{Tr} \left(\phi \wedge \partial\phi + \frac{2}{3} \phi \wedge \phi \wedge \phi \right) \quad (1.2)$$

with the trace taken in the adjoint representation of G . The F-term condition that follows from this action implies the vanishing of the $(2, 0)$ part of the Yang-Mills field strength so that (the lowest component of) ϕ describes an anti-holomorphic connection.

In this analogy, the vector multiplet we are ignoring plays the role of the gravitational fields we are freezing (with indices along X) and the chiral scalars ϕ stand in for the scalar fields in the tensor hierarchy (all indices along Y).

In the next section, we describe the embedding of the components of eleven-dimensional supergravity into superfields on $\mathbf{R}^{4|4} \times \mathbf{R}^7$. (The fields on $\mathbf{X} \times Y$ follow from this by covariantizing derivatives, as usual [12–14].) In section 3 we construct the action from the Chern-Simons invariant of the non-abelian tensor hierarchy and a supersymmetric extension of the Hitchin functional for Y . The equations of motion of the F- and D-auxiliary fields are computed. From this we obtain simultaneously the conditions for supersymmetry and the scalar potential. We conclude in section 5 with a discussion of our results. Appendix A contains a brief review of G_2 structures, the Hitchin functional, and some useful identities.

2 Superfields and Components

We begin by embedding the eleven-dimensional component fields into simple superspace. The eleven-dimensional supergravity component spectrum consists of a Riemannian metric g_{mn} or, more properly, its frame e_m^a , a 32-component Majorana gravitino ψ_m^α , and an abelian 3-form gauge field C_{mnp} . Here, bold indices are eleven-dimensional: $\alpha, \beta = 1, \dots, 32$ are Majorana spinor indices and we use the early-late convention for tangent vector indices $a, b = 0, \dots, 9$ and coordinate indices $m, n = 0, \dots, 9$. The bosonic part of the eleven-dimensional supergravity action is given by

$$\kappa^2 S_{11} = \int d^{11}x \sqrt{-g} \left[\frac{1}{2} R(g) - \frac{1}{4} F^2 \right] - \frac{1}{12} \int C \wedge F \wedge F. \quad (2.1)$$

Here, R is the Ricci scalar of the metric, g is its determinant, and $F = dC$ is the 4-form field strength of the gauge 3-form C .

As we will be embedding into a superspace modeled on $\mathbf{R}^{4|4} \times \mathbf{R}^7$, we must first reduce these components to “spacetime” X and “internal” Y :

$$e_m^a, \quad e_m^i, \quad g_{ij}, \quad \psi_m^{\alpha I}, \quad \psi_i^{\alpha I}, \quad C_{mnp}, \quad C_{mni}, \quad C_{mij}, \quad C_{ijk}. \quad (2.2)$$

The new indices on $X \times Y$ are as follows: $m, n = 0, \dots, 3$ denote spacetime coordinate indices, $a, b = 0, \dots, 3$ are spacetime tangent vectors indices, $i, j = 1, \dots, 7$ will be taken to be internal coordinate indices, $\alpha, \dot{\alpha} = 1, 2$ are $SL(2, \mathbf{C})$ indices, and finally, $I, J = 1, \dots, 8$ stand for $SO(8)$ R-symmetry indices.

To embed in superfields of $\mathbf{R}^{4|4} \times \mathbf{R}^7$, it is necessary to split up the gravitino fields and put one of them into an irreducible superspin- $\frac{3}{2}$ multiplet with the frame e_m^a . This will then be the 4D, $N = 1$ super-frame E_M^A . The other 7 gravitini must then go into a superspin-1 multiplet $\Psi^{\alpha i}$ transforming in the defining representation of the $SO(7) \subset SO(8)_R$ subgroup. (For notational simplicity, we do not distinguish between coordinate indices on Y and this subgroup.) The remaining fields consist of 1 3-form, 7 2-forms, $21 + 7 = 28$ vectors, and $28 + 35 = 63$ scalars (and their spin- $\frac{1}{2}$ superpartners). This set of fields is encoded in a non-abelian tensor hierarchy [6, 7] as we review presently.

Since there are many (super)fields involved, we try to minimize notation as follows: For any superfield X we define supersymmetry-covariant descendant superfields by acting with superspace derivatives. The descendants with the same statistics as X are defined by

$$\begin{aligned} f_X &= -\frac{1}{4}\bar{D}^2 X, \quad A_{Xa} = -\frac{1}{4}(\tilde{\sigma}_a)^{\dot{\alpha}\alpha}[D_\alpha, \bar{D}_{\dot{\alpha}}]X, \\ \tilde{f}_X &= -\frac{1}{4}D^2 X, \quad d_X = \frac{1}{32}\{D^2, \bar{D}^2\}X, \end{aligned} \quad (2.3)$$

whereas those of opposite statistics are

$$\begin{aligned} \chi_{X\alpha} &= D_\alpha X, \quad w_{X\alpha} = -\frac{1}{4}\bar{D}^2 D_\alpha X, \\ \tilde{\chi}_{X\dot{\alpha}} &= \bar{D}_{\dot{\alpha}} X, \quad \tilde{w}_{X\dot{\alpha}} = -\frac{1}{4}D^2 \bar{D}_{\dot{\alpha}} X. \end{aligned} \quad (2.4)$$

These superfields are used to define the covariant component fields by projecting $\theta, \bar{\theta} \rightarrow 0$ an operation we denote with a “|”. In terms of these components, the superfield can be written as

$$\begin{aligned} X &= X| + \theta^\alpha \chi_{X\alpha}| + \bar{\theta}_{\dot{\alpha}} \tilde{\chi}_{X\dot{\alpha}}| + \theta^2 f_X| + \bar{\theta}^2 \tilde{f}_X| - \theta \sigma^a \bar{\theta} A_{Xa}| \\ &\quad + \bar{\theta}^2 \theta^\alpha \left(w_{X\alpha}| + \frac{i}{2} \sigma_{\alpha\dot{\alpha}}^a \partial_a \tilde{\chi}_{X\dot{\alpha}}| \right) + \theta^2 \bar{\theta}_{\dot{\alpha}} \left(\tilde{w}_{X\dot{\alpha}}| + \frac{i}{2} \tilde{\sigma}^{a\dot{\alpha}\alpha} \partial_a \chi_{X\alpha}| \right) + \theta^2 \bar{\theta}^2 \left(d_X| - \frac{1}{4} \square X| \right). \end{aligned} \quad (2.5)$$

Henceforth, we will drop the “|” notation on the right-hand side of such expansions. When X is real, the tilded fields are conjugate to the untilded ones and X , A_X , and d_X are real. When X is chiral, the tilded fields are absent and the remaining components are complex.

In this work, we will not have much need for the superspin- s fields with $s = \frac{3}{2}$ (superframe E_M^A) and 1 (seven gravitino superfields $\Psi^{\alpha i}$) so we will be brief. (For an explicit construction of the quadratic action of 5D, $N = 1$ supergravity analog in terms of these superfields, see refs. [3, 4].) At the linearized level, the conformal graviton can be described by the real superfield

$$H^a = \cdots + \theta \sigma^m \bar{\theta} e_m^a + \bar{\theta}^2 (\sigma^{ab} \theta)_\alpha \psi_b^\alpha + \theta^2 (\tilde{\sigma}^{ab} \bar{\theta})_{\dot{\alpha}} \bar{\psi}_b^{\dot{\alpha}} + \theta^2 \bar{\theta}^2 d^a. \quad (2.6)$$

It contains the (linearized) frame, the $N = 1$ gravitino, and a real auxiliary vector field. (Here and in the following ellipses will stand for components that can be removed by a choice of Wess-Zumino gauge.) The gravitino superfield

$$\Psi^{\alpha i} = \dots + (\sigma^a \bar{\theta})^\alpha B_a{}^i + (\theta \sigma^m \bar{\theta}) \psi_m^{\alpha i} + \bar{\theta}^2 \theta^\alpha w^i + \theta^2 (\sigma^a \bar{\theta})^\alpha y_a^i + \bar{\theta}^2 (\theta \sigma^{ab})^\alpha w_{ab}^i + \dots \quad (2.7)$$

carries the 7 remaining gravitini, 7 “graviphotons”, and a collection of auxiliary fields, the precise content of which depends on the structure of the supergravity gauge transformation (which we will not need).

All but one of the remaining bosonic fields can be embedded in an abelian tensor hierarchy of superfields [6]. This is a chain complex of superfields constructed by combining the superspace analog of the de Rham complex on X and the de Rham complex on Y . The components of the 11D 3-form C_{abc} fit into the elements of this complex as follows:

$$\Phi_{ijk} = C_{ijk} + iF_{ijk} + \dots + \theta^2 f_{ijk} \quad (2.8a)$$

$$V_{ij} = \dots + \theta \sigma^a \bar{\theta} C_{a ij} + \dots + \theta^2 \bar{\theta}^2 d_{ij} \quad (2.8b)$$

$$\Sigma_{\alpha i} = \dots + \theta_\alpha H_i + (\theta \sigma^{ab})_\alpha C_{abi} + \dots \quad (2.8c)$$

$$X = \dots + \bar{\theta}^2 G + \theta^2 \bar{G} + \theta \sigma_a \bar{\theta} \epsilon^{abcd} C_{bcd} + \dots + \theta^2 \bar{\theta}^2 d_X. \quad (2.8d)$$

General abelian Chern-Simons-like invariants of this hierarchy in superspace were constructed in [6]. For eleven-dimensional supergravity, this is a cubic invariant of this superspace complex.

Being valued in the exterior algebra of the internal space, these fields are “charged” under the mixed components of the frame gauging the $\mathfrak{g} = \mathfrak{diff}(Y)$ symmetry. This results in a non-abelian gauging of the tensor hierarchy by a super- \mathfrak{g} -connection with spinorial superfield

$$\mathcal{A}_\alpha^i = \dots + (\sigma^a \bar{\theta})^\alpha e_a{}^i + \dots + \theta^2 \bar{\theta}^2 \mathbf{d}^i \quad (2.9)$$

arising by gauge-covariantizing the flat superspace derivative $D_\alpha \rightarrow \mathcal{D}_\alpha$ (minimal coupling). At this point the non-abelian tensor hierarchy is a \mathfrak{g} -equivariant super-de Rham complex of forms on $\mathbf{X} \times Y$. Its Chern-Simons-like invariant was studied in some detail and generality in [7]. We review the gauge transformations, \mathfrak{g} -covariant superfield strengths, Bianchi identities, and Chern-Simons action in section 3.1.

It remains to discuss the fate of the 28 metric scalars g_{ij} . Although we have not embedded them explicitly, we expect that they can be accounted for by the real scalars F_{ijk}

in the chiral field of the tensor hierarchy. (We elaborate on this in section 3.2.) Assuming this, we have embedded the component fields (2.2) of eleven-dimensional supergravity into a collection of prepotentials consisting of the conformal supergraviton H^a (2.6), 7 gravitino superfields $\Psi^{\alpha i}$ (2.7), and the superfields of the gauged tensor hierarchy (2.8 and 2.9). As we will be freezing the conformal graviton and gravitino superfields, the remaining set of auxiliary fields come only from the tensor hierarchy. They consist of the components

$$d_X, \quad \mathbf{d}^i = d_{\mathcal{V}^i}, \quad d_{ij} = d_{V_{ij}}, \quad f_{ijk} = f_{\Phi_{ijk}}. \quad (2.10)$$

(Here \mathcal{V}^i is the prepotential of the non-abelian gauge field $\mathcal{A}_\alpha^i \partial_i \sim e^{-i\mathcal{V}\partial}(D_\alpha e^{i\mathcal{V}\partial})$.) In the next section, we will propose an action constructed from the superfields of this section and project to components, focusing on this set of auxiliary fields.

3 Action

In this section we propose an action constructed from the supergravity and tensor hierarchy fields to lowest order in a gravitino superfield expansion (1.1). This is a superspace action consisting of a Chern-Simons term and a generalization of the Hitchin functional.

3.1 Chern-Simons Action

The superspace Chern-Simons action appropriate to eleven-dimensional supergravity is the cubic invariant of the non-abelian tensor hierarchy describing a gauge 3-form. The embedding of the components of this 3-form into superfields is represented in equation (2.8). The gauge transformations for these prepotentials are²

$$\delta\Phi = \mathcal{L}_\lambda \Phi + \partial\Lambda \quad (3.1a)$$

$$\delta V = \mathcal{L}_\lambda V + \frac{1}{2i}(\Lambda - \bar{\Lambda}) - \partial U \quad (3.1b)$$

$$\delta\Sigma_\alpha = \mathcal{L}_\lambda \Sigma_\alpha - \frac{1}{4}\bar{\mathcal{D}}^2 \mathcal{D}_\alpha U + \partial\Upsilon_\alpha + \iota_{\mathcal{W}_\alpha} \Lambda \quad (3.1c)$$

$$\delta X = \mathcal{L}_\lambda X + \frac{1}{2i} \left(\mathcal{D}^\alpha \Upsilon_\alpha - \bar{\mathcal{D}}_{\dot{\alpha}} \bar{\Upsilon}^{\dot{\alpha}} \right) - \omega_h(\mathcal{W}_\alpha, U). \quad (3.1d)$$

²We attempt to give a self-contained description of that part of the work on Chern-Simons-like invariants of the non-abelian tensor hierarchy needed for this paper, but please see references [6, 7] for additional material and information on this important ingredient of the construction.

All fields are differential forms on Y ; ∂ denotes the de Rham differential and wedge products are implied. The abelian part of the gauge transformation is parameterized by the superfields Λ_{ij} (chiral), U_i (real), and Υ_α (chiral) encoding the components of an eleven-dimensional super-2-form. The non-abelian part $\mathfrak{g} = \mathfrak{diff}(Y)$ acts by the Lie derivative with respect to the real scalar superfield λ^i . To check gauge invariance, we must use separation of the Lie derivative into the de Rham differential and the contraction operator ι using Cartan's formula $\mathcal{L}_V = \partial\iota_V + \iota_V\partial$. The composite superfield ω_h is the so-called ‘‘Chern-Simons superform’’. For any chiral spinor superfield χ_α and real scalar superfield v ,

$$\begin{aligned}\omega_h(\chi_\alpha, v) &:= \iota_{\chi^\alpha} \mathcal{D}_\alpha v + \iota_{\bar{\chi}_{\dot{\alpha}}} \bar{\mathcal{D}}^{\dot{\alpha}} v + \frac{1}{2} \left(\iota_{\mathcal{D}^\alpha \chi_\alpha} v + \iota_{\bar{\mathcal{D}}_{\dot{\alpha}} \bar{\chi}^{\dot{\alpha}}} v \right) \\ \Rightarrow \quad \bar{\mathcal{D}}^2 \omega_h(\chi_\alpha, v) &= \iota_{\chi^\alpha} \bar{\mathcal{D}}^2 \mathcal{D}_\alpha v + \frac{1}{2} \bar{\mathcal{D}}^2 \iota_{\mathcal{D}^\alpha \chi_\alpha - \bar{\mathcal{D}}_{\dot{\alpha}} \bar{\chi}^{\dot{\alpha}}} v.\end{aligned}\tag{3.2}$$

(Its name derives from the fact that if $\chi \sim \bar{D}^2 Dv$ is the field strength superfield of the real vector superfield v , then the second term vanishes and $\bar{D}^2 \omega \sim \chi^2$ gives the superspace analog of $d\omega = F \wedge F$.)

The non-abelian gauge field strength $\mathcal{W}^{\alpha i}$ is defined by $[\mathcal{D}_a, \bar{\mathcal{D}}_{\dot{a}}] = (\sigma_a)_{\alpha\dot{\alpha}} \mathcal{L}_{\mathcal{W}^\alpha}$. The field strengths $\partial\Phi$ and

$$F = \frac{1}{2i} (\Phi - \bar{\Phi}) - \partial V \tag{3.3a}$$

$$W_\alpha = -\frac{1}{4} \bar{\mathcal{D}}^2 \mathcal{D}_\alpha V + \partial \Sigma_\alpha + \iota_{\mathcal{W}_\alpha} \Phi \tag{3.3b}$$

$$H = \frac{1}{2i} \left(\mathcal{D}^\alpha \Sigma_\alpha - \bar{\mathcal{D}}_{\dot{\alpha}} \bar{\Sigma}^{\dot{\alpha}} \right) - \partial X - \omega_h(\mathcal{W}_\alpha, V) \tag{3.3c}$$

$$G = -\frac{1}{4} \bar{\mathcal{D}}^2 X + \iota_{\mathcal{W}^\alpha} \Sigma_\alpha \tag{3.3d}$$

are invariant under the abelian transformations and covariant under the non-abelian ones: $\delta(FS) = \mathcal{L}_\lambda(FS)$. Being given explicitly in terms of the prepotential superfields, these field strengths identically satisfy the Bianchi identities

$$\frac{1}{2i} (\partial\Phi - \partial\bar{\Phi}) = \partial F \tag{3.4a}$$

$$-\frac{1}{4} \bar{\mathcal{D}}^2 \mathcal{D}_\alpha F = -\partial W_\alpha - \iota_{\mathcal{W}_\alpha} \partial\Phi \tag{3.4b}$$

$$\frac{1}{2i} \left(\mathcal{D}^\alpha W_\alpha - \bar{\mathcal{D}}_{\dot{\alpha}} \bar{W}^{\dot{\alpha}} \right) = \partial H + \omega_h(\mathcal{W}_\alpha, F) \tag{3.4c}$$

$$-\frac{1}{4} \bar{\mathcal{D}}^2 H = -\partial G - \iota_{\mathcal{W}^\alpha} W_\alpha \tag{3.4d}$$

$$\bar{\mathcal{D}}_{\dot{\alpha}} G = 0. \tag{3.4e}$$

(Equivalently, 3.3 is the solution to these constraints.) In terms of these prepotentials and field strength superfields, the Chern-Simons super-invariant $S_{CS} = \int d^4x \int_Y L_{CS}$ is

defined by the Lagrangian

$$\begin{aligned}
-12\kappa^2 L_{CS} = & i \int d^2\theta \Phi \wedge [\partial\Phi G + \tfrac{1}{2}W^\alpha \wedge W_\alpha - \tfrac{i}{4}\bar{\mathcal{D}}^2(F \wedge H)] \\
& + \int d^4\theta V \wedge [\partial\Phi \wedge H + F \wedge \mathcal{D}^\alpha W_\alpha + 2\mathcal{D}^\alpha F \wedge (W_\alpha - i\iota_{W_\alpha}F)] \\
& + i \int d^2\theta \Sigma^\alpha \wedge [\partial\Phi \wedge W_\alpha - \tfrac{i}{4}\bar{\mathcal{D}}^2(F \wedge \mathcal{D}_\alpha F)] \\
& - \int d^4\theta X \partial\Phi \wedge F + \text{h.c.}
\end{aligned} \tag{3.5}$$

Using the Bianchi identities, one can check that this Lagrangian transforms into an exact superform under the gauge transformations (3.1). Comparing the first term with the superpotential $\int d^2\theta \Phi \partial\Phi$ put forward in reference [15] suggests an interpretation of this action as the covariantization of the G_2 superpotential under the tensor hierarchy transformations.

The tensor hierarchy can be coupled to gravity by replacing $\mathcal{D} \rightarrow \mathcal{D}$ with the gravitationally covariant superspace derivative, covariantizing the measures as usual, and replacing $\bar{\mathcal{D}}^2 \rightarrow \bar{\mathcal{D}}^2 - 8R$ [16, 17]. (Equivalently, one can replace \mathcal{D} by the conformal superspace supergravity derivative [21, 22].) Then, the complete component projection of the action can be computed straightforwardly, but for this paper we interested only in the contribution to the scalar potential of the component theory. This simplifies the calculations significantly.

Firstly, we can ignore the supergravity couplings so the action in the form (3.5) will suffice to compute the potential. Next, the potential consists of only internal derivatives of the scalars g_{ij} and C_{ijk} in the hierarchy so we will drop all gauge fields (with X indices) and spacetime derivatives. (Of course there are internal derivatives of gauge fields but we find it more convenient to think of them as covariantizing the spacetime derivatives.) The gauge superfields can still contribute F- and D-type auxiliary fields so we will keep those. An exception is the superfield $\Sigma_{\alpha i}$ for the gauge 2-form C_{abi} : This representation (2.8c) has no auxiliary fields so we will remove it altogether. (We revisit the validity of this simplification in section 5.) Performing the Graßmann integration, and focusing only on the remaining fields, we find

$$\begin{aligned}
\kappa^2 L_{CS} = & \frac{1}{288} \epsilon^{ijklmnp} \left[F_{ijkl} F_{mnp} d_X - 4F_{ijk} \mathbf{d}^r F_{lmr} \mathbf{d}^s F_{nps} + 12F_{ijk} d_{lm} d_{np} \right. \\
& \left. - \frac{1}{2} \left(G[4F_{ijk} \partial_l f_{mnp} + iF_{ijkl} f_{mnp}] + \text{h.c.} \right) \right]. \tag{3.6}
\end{aligned}$$

Here $F_{ijkl} = 4\partial_{[i}C_{jkl]}$ is the lowest component of the real part of $\partial\Phi$.³ We use this result in section 4 once we have constructed the remaining terms to which we turn next.

3.2 Kähler Action

In the previous section, we reviewed the Chern-Simons action arising in the gravitino superfield expansion (1.1). The remaining terms at this order in the expansion define the Kähler action S_K . In this section, we propose an explicit formula for this action.

In the component embedding of section 2, we assumed the metric scalars (pull-back of the 11D metric to Y) appeared in the superfield spectrum in the imaginary part of the 35 chiral fields carrying the 3-form scalars (the real part). In section 3.1, we reviewed the construction of the superfield strength F_{ijk} associated to these chiral fields (eq. 3.3a). This field strength is invariant under the abelian 3-form transformations as should be the metric scalars. Therefore, we identify the dynamical metric scalars as the fluctuations of this field strength and construct a Riemannian metric following Hitchin: First, define the superfield

$$s_{ij}(F) := -\frac{1}{144}\epsilon^{klmnpqr}F_{ikl}F_{jmn}F_{pqr}. \quad (3.7)$$

Stability of F means that $\det(s) \neq 0$. Setting $g_{ij}(F) := \det^{-1/9}(s)s_{ij}$, defines a Riemannian metric on Y . We note that the replacement $\Phi_{ijk} \rightarrow F_{ijk}$ guarantees abelian gauge invariance but it is not holomorphic. This has the important consequence that we cannot use $g_{ij}(F)$ to construct invariant superpotential terms.

Taking the determinant, we can define the superspace analog of the Riemannian volume on Y . So motivated, we propose to take as the Kähler action the following natural generalization of the superspace volume of $\mathbf{X} \times Y$:

$$S_K = -\frac{3}{\kappa^2} \int d^4x \int d^7y \int d^4\theta E \sqrt{g(F)} \left[(\bar{G}G)^{1/3} - \frac{1}{3}(\partial_i H_a)^2 \right]. \quad (3.8)$$

Here, $E = \text{sdet}(E_M^A)$ is the super-determinant of the 4D, $N = 1$ components of the frame. The second term is the $N = 1$ super-graviton “mass” term. (This term is needed because there are no spin-2 auxiliary fields.) Its normalization is fixed by eleven-dimensional Lorentz invariance: In the quadratic approximation, $-3 \int d^4x d^7y d^4\theta E \rightarrow -\int d^4x d^7y H_a \square H^a$ in the gauge $\bar{D}_\alpha H^a = 0$ [3, 12, 14].

³This should not be confused with F_{ijk} which denotes the lowest component of the field strength (3.3a). (G is the lowest component of (3.3d), and the remaining auxiliary fields are defined in (2.10).)

The first term looks like the old-minimal supergravity action [12, 14], provided we identify $G \leftrightarrow \Phi_0$ with the chiral conformal compensator.⁴ We will revisit this connection in section 5 but if we simply assume it for now, then freezing gravity amounts to setting

$$H^a \rightarrow \theta \sigma^a \bar{\theta} \quad , \quad \Psi^{\alpha i} \rightarrow 0 \quad , \quad \text{and} \quad G \rightarrow 1 + \theta^2 (d_X - \frac{i}{4!} \epsilon^{abcd} F_{abcd}). \quad (3.9)$$

As we are interested only in the scalar potential in this paper, we drop the field strength of the component 3-form C_{abc} . (However, matching the coefficient of this component to the correct value fixes the G -dependence in eq. 3.8. In the setting in which 4-form fluxes are turned on, these terms give quantum corrections to the potential as explained by Beasley and Witten [19].) This reduces the Kähler part of the action to

$$S_K \rightarrow -\frac{3}{\kappa^2} \int d^4x \int d^7y \int d^4\theta \sqrt{g(F)} \left[1 + \frac{1}{3}(\theta^2 + \bar{\theta}^2)d_X + \frac{1}{9}\theta^2\bar{\theta}^2 d_X^2 \right]. \quad (3.10)$$

We recognize the leading term as a superspace version of the Hitchin functional (A.3) for the G_2 structure on Y [8, 9] (see also [20]).

Integrating over the odd coordinates and collecting the auxiliary field terms needed to compute the scalar potential, we find

$$\begin{aligned} \kappa^2 L_K = & -\frac{1}{3}\sqrt{g}d_X^2 - \frac{1}{18}\sqrt{g}F^{ijk}\text{Im}(f_{ijk})d_X + \frac{1}{6}\sqrt{g}F^{ijk}F_{ijkl}\mathbf{d}^l \\ & - \frac{1}{2}\sqrt{g}F^{ijk}\partial_k d_{ij} + \frac{3}{4}G^{ijk,mnp}\bar{f}_{ijk}f_{mnp}. \end{aligned} \quad (3.11)$$

Here,

$$G^{ijk,mnp} := \frac{1}{3!3!}\sqrt{g}g^{[i|m}g^{|j|n}g^{|k|p]} + \frac{1}{18\cdot 3!\cdot 3!}\sqrt{g}F^{ijk}F^{mnp} + \frac{1}{4!}\sqrt{g}g^{[m|[i}\psi^{jk]|np]} \quad (3.12)$$

is essentially the Hitchin metric on the moduli space of (complexified) G_2 structures [8]. In terms of G_2 projections

$$18\omega_{ijk}G^{ijk,lmn}\omega_{lmn} = -\frac{4}{3}\omega_{1ijk}^2 - \omega_{7ijk}^2 + \omega_{27ijk}^2 \quad (3.13)$$

for any 3-form ω .

⁴Actually, this would be a slightly modified version of old-minimal supergravity since G has a real prepotential X (3.3d and 2.8d). Such a modification of the compensator was first exploited in [18] to simplify supergraph calculations. It was also needed in the (4+1)-dimensional version of the superspace supergravity considered here [3–5].

4 Scalar Potential

Using various identities from G_2 linear algebra collected in appendix A, we can rewrite the Chern-Simons contribution (3.6) as

$$\begin{aligned} \kappa^2 L_{CS} = & 2\sqrt{g}g_{ij}\mathbf{d}^i\mathbf{d}^j + \frac{1}{4}\sqrt{g}\psi^{ijkl}d_{ij}d_{kl} + \frac{1}{48}\sqrt{g}\psi^{ijkl}F_{ijkl}d_X \\ & + \frac{1}{12}\partial_i(\sqrt{g}\psi^{ijkl})\text{Re}(f_{jkl}) + \frac{1}{288}\epsilon^{ijklmnp}\left[F_{ijkl}\text{Im}(f_{mnp}) + 4\partial_i F_{jkl}\text{Re}(f_{mnp})\right]. \end{aligned} \quad (4.1)$$

To proceed, it is useful to introduce the intrinsic torsion forms τ_μ for $\mu = 0, 1, 2, 3$ and analogous quantities σ_μ for $\mu = 0, 1, 3$ defined by [10]

$$d\varphi = \tau_0\psi + 3\tau_1\varphi + *\tau_3, \quad d\psi = 4\tau_1\psi + \tau_2\varphi, \quad (4.2a)$$

$$dC = \sigma_0\psi + 3\sigma_1\varphi + *\sigma_3. \quad (4.2b)$$

(We could make the analogous definition for the the components of $d*C$ but the action depends only on C and dC ; the C -field analogue of the torsion class τ_2 is not gauge invariant.)

In terms of these intrinsic torsion forms, $\frac{1}{\sqrt{g}}\partial_k(\sqrt{g}F^{ijk}) = 4F^{ijk}(\tau_1)_k + (\tau_2)^{ij}$ which we use to rewrite (3.11) as

$$\begin{aligned} \kappa^2 L_K = & -\frac{1}{3}\sqrt{g}d_X^2 - \frac{1}{18}\sqrt{g}F^{ijk}\text{Im}(f_{ijk})d_X + 2\sqrt{g}d^{ij}\left[F_{ijk}(\tau_1)^k + \frac{1}{4}(\tau_2)_{ij}\right] \\ & + 12\mathbf{d}^i(\sigma_1)_i + \frac{3}{4}\bar{f}_{ijk}G^{ijk,mnp}f_{mnp}. \end{aligned} \quad (4.3)$$

In terms of G_2 representations this becomes

$$\begin{aligned} \kappa^2 L_K = & -\frac{1}{3}\sqrt{g}d_X^2 - \frac{1}{18}\sqrt{g}F^{ijk}\text{Im}(f_{ijk})d_X + 2\sqrt{g}d^{ij}\left[F_{ijk}(\tau_1)^k + \frac{1}{4}(\tau_2)_{ij}\right] \\ & + 12\mathbf{d}^i(\sigma_1)_i - \frac{1}{18}|f_{1ijk}|^2 - \frac{1}{24}|f_{\mathbf{2}ijk}|^2 + \frac{1}{24}|f_{\mathbf{2}\mathbf{7}ijk}|^2. \end{aligned} \quad (4.4)$$

Combining this result with (4.1), we find for the equations of motion of the auxiliary fields

$$d_X = \frac{21}{4}\sigma_0 - \frac{1}{12}F^{ijk}\text{Im}(f_{ijk}) \quad (4.5a)$$

$$g_{ij}\mathbf{d}^j = -3(\sigma_1)_i \quad (4.5b)$$

$$\psi_{ijkl}d^{kl} = -\left[4F_{ijk}g^{kl}(\tau_1)_l + (\tau_2)_{ij}\right] \quad (4.5c)$$

$$G^{ijk,lmn}\bar{f}_{lmn} = \frac{i}{432}\epsilon^{ijklmnp}\bar{E}_{lmnp} + \frac{1}{18}\partial_l\psi^{ijkl} - \frac{i}{27}\sqrt{g}d_X F^{ijk} \quad (4.5d)$$

To solve these, it is convenient to decompose into G_2 representations and invert on irreducible representations using the identities in appendix A. Doing so gives

$$\begin{aligned} d_X = & -\frac{7}{2}\sigma_0, \quad \mathbf{d}^i = -3(\sigma_1)^i, \quad d_{\mathbf{7}ij} = F_{ijk}(\tau_1)^k, \quad d_{\mathbf{14}ij} = -\frac{1}{2}(\tau_2)_{ij}, \\ \bar{f}_{\mathbf{1}ijk} = & -\left(\frac{3}{4}\tau_0 + \frac{5}{2}i\sigma_0\right)F_{ijk}, \quad \bar{f}_{\mathbf{7}ijk} = -3\psi_{ijkl}(\tau_1 + i\sigma_1)^l, \quad \bar{f}_{\mathbf{2}\mathbf{7}ijk} = (\tau_3 + i\sigma_3)_{ijk}. \end{aligned} \quad (4.6)$$

Note that the F- and D-flatness conditions are equivalent to the vanishing of each of the intrinsic torsion forms ($\tau_\mu = 0$) and their gauge field counterparts ($\sigma_\mu = 0$ for $\mu \neq 2$).

In terms of G_2 projections, we may write the potential in terms of auxiliary fields as

$$V = \frac{1}{4}d_X^2 + 2g_{ij}\mathbf{d}^i\mathbf{d}^j - d_{7ij}^2 + \frac{1}{2}d_{14ij}^2 - \frac{1}{18}|f_{1ijk} + \frac{i}{2}d_X F_{ijk}|^2 - \frac{1}{24}|f_{7ijk}|^2 + \frac{1}{24}|f_{27ijk}|^2. \quad (4.7)$$

Here we used (3.13) and the fact that for any 2-form η ,

$$\eta_{ij}\psi^{ijkl}\eta_{kl} = -4\eta_{7ij}^2 + 2\eta_{14ij}^2 \quad (4.8)$$

as follows from (A.8). Substituting the algebraic equations (4.6) back into the component action and using (A.6), we find the following scalar potential:

$$\begin{aligned} V(x, y) &= -\frac{42}{18} \cdot \frac{9}{16}\tau_0^2 - \left(6 + \frac{9 \cdot 24}{24}\right)\tau_1^2 + \frac{1}{2} \cdot \frac{1}{4}\tau_2^2 + \frac{1}{24}\tau_3^2 \\ &\quad + \left(\frac{49}{16} - \frac{42}{18} \cdot \frac{9}{16}\right)\sigma_0^2 + \left(2 \cdot 9 - \frac{9 \cdot 24}{24}\right)\sigma_1^2 + \frac{1}{24}\sigma_3^2 \\ &= -\frac{21}{16}\tau_0^2 - 15\tau_1^2 + \frac{1}{8}\tau_2^2 + \frac{1}{24}\tau_3^2 + \frac{7}{4}\sigma_0^2 + 9\sigma_1^2 + \frac{1}{24}\sigma_3^2 \end{aligned} \quad (4.9)$$

To interpret this, we appeal to a result of Bryant who has computed that the scalar curvature of the metric is given by [10]

$$R(g) = 12\nabla^i(\tau_1)_i + \frac{21}{8}\tau_0^2 + 30\tau_1^2 - \frac{1}{4}\tau_2^2 - \frac{1}{12}\tau_3^2 \quad (4.10)$$

when written in terms of the intrinsic torsion forms (4.2a). Similarly, one checks that the σ_μ terms combine into a Maxwell term. Therefore, up to a surface term, we have shown that

$$-V(x, y) = \frac{1}{2}R(g) - \frac{1}{4 \cdot 4!}F_{ijkl}^2, \quad (4.11)$$

which is the potential of the bosonic part of eleven-dimensional supergravity. We conclude that the action (1.1) reproduces the correct scalar potential of eleven-dimensional supergravity.

5 Discussion

In this paper we computed the potential of eleven-dimensional supergravity as it is described in a superspace background of the form $\mathbf{R}^{4|4} \times Y$ with Y a (not necessarily compact) manifold with G_2 structure. To first order in a gravitino superfield expansion (1.1), the action is the sum of two terms. The first is the superspace Chern-Simons invariant

(3.5) of the gauged tensor hierarchy of the eleven-dimensional 3-form. This tensor hierarchy is a \mathfrak{g} -equivariant superspace chain complex with \mathfrak{g} the algebra of diffeomorphisms on Y [6, 7].

The second is a superspace version of the Hitchin functional for G_2 -structure manifolds (3.8). The bosonic version of this functional is the volume of Y as computed from the Riemannian metric constructed from an arbitrary (stable) 3-form. (The stationary points of this functional on cohomology classes of the 3-form define G_2 -holonomy metrics.) This functional is lifted to superspace by formally replacing the 3-form with the tensor hierarchy superfield strength containing the gravitational scalars and integrating over superspace. (The 3-form on Y on which the tensor hierarchy is based is embedded as the imaginary part of a chiral superfield; the stable 3-form is the real part.)

Having defined the action thus, we computed its potential by integrating out the auxiliary fields of the multiplets in the tensor hierarchy. What we find is the Einstein-Hilbert action on Y in the form computed by Bryant [10] with an analogous form for the Maxwell term for the 3-form scalars. As this is the correct potential for eleven-dimensional supergravity on $X \times Y$, this observation relates topological M theory [20] to “physical” M-theory, and suggests the following construction of eleven-dimensional supergravity on $\mathbf{X} \times Y$ to this order in the super-gravitino expansion (1.1): Starting with the non-abelian tensor hierarchy, one defines on the space of 3-form field strength superfields the curved superspace generalization (3.8) of the Hitchin functional. To this one adds the Chern-Simons super-invariant (3.5) in curved superspace [16] (see also [21, 22]). Note that we cannot add a superpotential beyond the F-terms coming from the Chern-Simons action because the G_2 -structure metric is not chiral. Therefore, we might expect this to be the full answer at this order of the gravitino expansion.

As it stands, this proposal probably requires some modification, or at least a better understanding of the following puzzle: In freezing the supergravity fields (3.9) and ignoring the superfields containing the gauge 2-forms, we are implicitly assuming that these fields do not carry propagating scalars. In particular, one needs to explain the mechanism by which the superfluous scalars in the 3- and 2-form multiplets are removed from the spectrum.

A potential resolution to this problem is that the 4-form field strength actually *is* the supergravity conformal compensator $G = \Phi_0^3$. Similarly the 3-form field strengths H_i would be the $Spin(7)/G_2$ compensators. In fact, the Hitchin metric is negative definite on these representations (cf. eq. 3.13) which would result in the wrong-sign kinetic terms for these fields that is the hallmark of a compensating field [12]. In such a scenario the super-diffeomorphisms of the theory would allow one to fix G and H_i by a choice of

conformal and $SO(7)$ gauge.

In this interpretation, the $\int G\Phi\partial\Phi$ term in the Chern-Simons action (3.5) becomes

$$-\frac{i}{12\kappa^2} \int d^4x \int_Y \int d^2\theta \Phi_0^3 \Phi \wedge \partial\Phi + \text{h.c.} \quad (5.1)$$

This is the gravitational covariantization of the chiral superpotential postulated in reference [15] in the old-minimal formulation of 4D, $N = 1$ supergravity [23, 24]. This superpotential combines the action constructed by Gukov [25] (see also [26]) in the context of flux compactifications with the terms needed for holomorphicity in general backgrounds.

We can extend this interpretation to the Kähler term as follows by covariantizing of the Hitchin functional [12]

$$\int d^4x \int d^7y \int d^4\theta E \bar{\Phi}_0 \Phi_0 \sqrt{g(F)}. \quad (5.2)$$

Under the identification of the compensator as the cube root of G , we recognize this as the (non-mass part of the) Kähler action (3.8).

As discussed in references [19, 27, 28], the original flux potential gets corrections from M2-brane domain walls localized at points on Y and/or M5-branes wrapped on associative 3-cycles in Y . There it is argued that these corrections change the superpotential by what is essentially the current of the Page charge [29]. They then write the corrected Gukov superpotential as $W \sim m \text{vol}(Y) + \int_Y \Phi \wedge d\Phi$ where m is the Freund-Rubin mass [30]. Going back to superspace, the contribution of the first term would come from an expression of the form (5.2).

Finally, we mention that without including the (spacetime scalar) auxiliary fields of all of the superfields in the theory, one does not expect to recover the correct scalar potential. Since 4D, $N = 1$ supergravity contains two such fields, these should have already been included in our analysis lest they over-correct the potential. In the identification above, these fields were already taken into account by identifying the supergravity scalar auxiliary with d_X and the pseudo-scalar with the dual of the 4-form field strength $F_{abcd} = 4\partial_{[a}C_{bcd]}$. An analogous resolution of the puzzle for the scalars in H_i will have to await (and hint at) the construction of the couplings to the gravitino multiplets. These couplings are currently under investigation.

Acknowledgements

It is a pleasure to thank Daniel Butter and Stephen Randall for enlightening discussions and Sergei Gukov and Nikita Nekrasov for comments and suggestions. WDL3 and DR are grateful to the Simons Center for Geometry and Physics for hospitality during the IX Simons Summer Workshop. This work is partially supported by NSF Focused Research Grant DMS-1159404 and the Mitchell Institute for Physics and Astronomy at Texas A&M University.

A G_2 Toolbox

In this appendix, we review the construction of the Hitchin functional for G_2 -structure 7-manifolds and collect various identities of G_2 linear algebra [10, 31, 32]. Let φ be a 3-form on Y and define the symmetric bilinear form

$$s_{ij} := -\frac{1}{144}\epsilon^{abcdefg}\varphi_{iab}\varphi_{cde}\varphi_{jfg}. \quad (\text{A.1})$$

The 3-form φ is “stable” iff $\det(s) \neq 0$.⁵ We assume this non-degeneracy condition throughout the paper. A stable 3-form on the tangent spaces of Y reduces the structure group $GL(7) \rightarrow G_2$. Thus, our assumption implies that Y is a G_2 -structure manifold.

Normalizing

$$g_{ij} = s^{-1/9}s_{ij} \quad \Leftrightarrow \quad \sqrt{g}g_{ij} = s_{ij}, \quad (\text{A.2})$$

defines the Riemannian metric g on Y . We can construct the Riemannian volume functional from the determinant of the metric

$$\Phi(\varphi) := \int_Y d^{11}y \sqrt{g(\varphi)} \quad (\text{A.3})$$

This expression is (equivalent to) the Hitchin functional on the space of stable 3-forms on Y [8]. In that reference, it is shown if (Y, φ) a closed G_2 -holonomy manifold, then φ is (closed by definition and) a critical point of Φ restricted to the cohomology class

⁵Stability as formulated in [8, 9] is in terms of open orbits of the $GL(n)$ action on the space of p -forms on the tangent bundle of an n -manifold Y . This condition is the precise criterion for when a volume form constructed from fractional powers of the p -form exists. In order for g_{ij} to be a good metric, we actually need that φ is positive, implying that s_{ij} and g_{ij} are positive definite, but this is only a slightly stronger condition, since if φ is stable then either φ or $-\varphi$ is positive. In this paper we will simply take stability to mean $\det(s) \neq 0$, and we will not always emphasize positivity.

$[\varphi] \in H^3(Y, \mathbf{R})$. Conversely, if φ is a critical point on a cohomology class of a closed oriented 7-manifold Y such that φ is stable, then φ defines a metric with G_2 holonomy. For any p -form ω , let $\omega_{\mathbf{i}} := \pi_{\mathbf{i}}\omega$ denote the projection to the \mathbf{i} -dimensional representation. Then, under a variation $\delta\varphi$ of the G_2 structure form

$$\delta g_{ij} = \varphi_{(i}{}^{kl} \left[\frac{1}{9}(\delta\varphi)_{\mathbf{1}} + \frac{1}{2}(\delta\varphi)_{\mathbf{27}} \right]_{j)kl}, \quad (\text{A.4})$$

the metric does not transform (to first order) under the $\mathbf{7}$ projection of the variation. We will not need these facts for this paper; we include them only to motivate the definition of the Hitchin functional. (To the interested reader, we recommend Karigiannis' thesis [32].)

We now review some G_2 linear algebra and define the projectors from representations of $SO(7)$ to those of G_2 . Under the reduction $SO(7) \rightarrow G_2$ of the structure group, the $\mathbf{21}$ -dimensional space of 2-forms on Y decomposes into G_2 representations as $\mathbf{21} = \mathbf{7} \oplus \mathbf{14}$. Similarly, the $\mathbf{35}$ -dimensional space of 3-forms on Y decomposes as $\mathbf{35} = \mathbf{1} \oplus \mathbf{7} \oplus \mathbf{27}$. (We review the explicit formulæ for the projectors to these representations presently.) We start by defining the dual $\psi = *\varphi$ with components

$$\psi_{ijkl} = \frac{1}{3!} \sqrt{g} \epsilon_{ijklmnp} g^{mm'} g^{nn'} g^{pp'} \varphi_{m'n'p'} \quad (\text{A.5})$$

(the opposite is $\varphi_{ijk} = \frac{1}{4!} \sqrt{g} \epsilon_{ijklmnp} g^{mm'} g^{nn'} g^{pp'} g^{qq'} \psi_{m'n'p'q'}$). Useful identities include

$$\begin{aligned} \varphi^{ijk} \varphi_{ij'k'} &= 2\delta_{[j}^j \delta_{k']}^k - \psi_{j'k'}{}^{jk} \quad , \quad \psi^{ijkl} \psi_{ijk'l'} = 8\delta_{[k}^k \delta_{l']}^l - 2\psi_{k'l'}{}^{kl} \quad , \\ \varphi^{ijk} \varphi_{ijk'} &= 6\delta_{k'}^k \quad , \quad \psi^{ijkl} \psi_{ijkl'} = 24\delta_{l'}^l \quad , \quad \varphi_i{}^{lm} \psi_{jklm} = -4\varphi_{ijk} \quad , \end{aligned} \quad (\text{A.6})$$

where indices are raised and lowered with the metric (A.2). These identities can be used to construct the projectors from the representations of $SO(7)$ onto the irreducible representations of G_2 : For any 2-form η and 3-form ω ,

$$\pi_7 \eta_{ij} = \left(\frac{1}{3} \delta_i^k \delta_j^l - \frac{1}{6} \psi_{ij}{}^{kl} \right) \eta_{kl} \quad (\text{A.7a})$$

$$\pi_{14} \eta_{ij} = \left(\frac{2}{3} \delta_i^k \delta_j^l + \frac{1}{6} \psi_{ij}{}^{kl} \right) \eta_{kl} \quad (\text{A.7b})$$

$$\pi_1 \omega_{ijk} = \frac{1}{42} \varphi_{ijk} \varphi^{i'j'k'} \omega_{i'j'k'} \quad (\text{A.7c})$$

$$\pi_7 \omega_{ijk} = \left(\frac{1}{4} \delta_i^{i'} \delta_j^{j'} \delta_k^{k'} - \frac{3}{8} \psi_{[ij}{}^{i'j'} \delta_{k]}^{k'} - \frac{1}{24} \varphi_{ijk} \varphi^{i'j'k'} \right) \omega_{i'j'k'} \quad (\text{A.7d})$$

$$\pi_{27} \omega_{ijk} = \left(\frac{3}{4} \delta_i^{i'} \delta_j^{j'} \delta_k^{k'} + \frac{3}{8} \psi_{[ij}{}^{i'j'} \delta_{k]}^{k'} + \frac{1}{56} \varphi_{ijk} \varphi^{i'j'k'} \right) \omega_{i'j'k'}. \quad (\text{A.7e})$$

Two useful identities on the space of 2-forms are

$$\psi_{ij}{}^{kl} \eta_{7kl} = -4\eta_{7kl} \quad , \quad \psi_{ij}{}^{kl} \eta_{14kl} = 2\eta_{14kl} \quad (\text{A.8})$$

Similarly, on the space of 3-forms,

$$\omega^2 := \omega^{ijk}\omega_{ijk} = \omega_1^2 + \omega_7^2 + \omega_{27}^2 \quad (\text{A.9a})$$

$$g^{ii'}\psi^{jj'k'}\omega_{ijk}\omega_{i'j'k'} = -4\omega_1^2 - 2\omega_7^2 + \frac{2}{3}\omega_{27}^2 \quad (\text{A.9b})$$

$$(\varphi^{ijk}\omega_{ijk})^2 = 42\omega_1^2. \quad (\text{A.9c})$$

References

- [1] Katrin Becker, Melanie Becker, and Daniel Robbins. Kaluza-Klein Theories Without Truncation. *JHEP*, 02:140, 2015. [[arXiv:1412.8198](#)].
- [2] W. Siegel and M. Roček. On Off-shell Supermultiplets. *Phys.Lett.*, B105:275, 1981. [[INSPIRE entry](#)].
- [3] William D. Linch III, Markus A. Luty, and J. Phillips. Five-dimensional supergravity in N=1 superspace. *Phys.Rev.*, D68:025008, 2003. [[hep-th/0209060](#)].
- [4] S. James Gates Jr., William D. Linch III, and J. Phillips. Field strengths of linearized 5-D, N=1 superfield supergravity on a three-brane. *JHEP*, 0502:036, 2005. [[hep-th/0311153](#)].
- [5] I.L. Buchbinder, S. James Gates Jr., Hock-Seng Goh, William D. Linch III, Markus A. Luty, Siew-Phang Ng, and J. Phillips. Supergravity loop contributions to brane world supersymmetry breaking. *Phys.Rev.*, D70:025008, 2004. [[hep-th/0305169](#)].
- [6] Katrin Becker, Melanie Becker, William D. Linch III, and Daniel Robbins. Abelian tensor hierarchy in 4D, N = 1 superspace. *JHEP*, 03:052, 2016. [[arXiv:1601.03066](#)].
- [7] Katrin Becker, Melanie Becker, William D. Linch III, and Daniel Robbins. Chern-Simons actions and their gaugings in 4D, N = 1 superspace. *JHEP*, 06:097, 2016. [http://\[arXiv:1603.07362\]](#).
- [8] Nigel J. Hitchin. The Geometry of Three-Forms in Six and Seven Dimensions. *J. Diff. Geom.*, 55(3):547–576, 2000. [[arXiv:math/0010054](#)].
- [9] Nigel J. Hitchin. Stable Forms and Special Metrics. 2001. [[arXiv:math/0107101](#)].
- [10] R. L. Bryant. Some remarks on G_2 -structures. *ArXiv Mathematics e-prints*, May 2003. [[arXiv:math/0305124](#)].
- [11] Neil Marcus, Augusto Sagnotti, and Warren Siegel. Ten-dimensional Supersymmetric Yang-Mills Theory in Terms of Four-dimensional Superfields. *Nucl. Phys.*, B224:159, 1983. [[INSPIRE entry](#)].
- [12] S.J. Gates, Marcus T. Grisaru, M. Roček, and W. Siegel. *Superspace Or One Thousand and One Lessons in Supersymmetry*. Frontiers in Physics, 58. Benjamin/Cummings, 1983. [[hep-th/0108200](#)].
- [13] J. Wess and J. Bagger. *Supersymmetry and supergravity*. Princeton, USA: Univ. Pr. (1992) 259 p, 1992. [[INSPIRE entry](#)].
- [14] I.L. Buchbinder and S.M. Kuzenko. *Ideas and methods of supersymmetry and supergravity: Or a walk through superspace*. Bristol, UK: IOP (1998) 656 p, 1998. [[INSPIRE entry](#)].
- [15] Katrin Becker, Daniel Robbins, and Edward Witten. The α' Expansion On A Compact Manifold Of Exceptional Holonomy. *JHEP*, 06:051, 2014. [[arXiv:1404.2460](#)].

- [16] Stephen Randall. Supersymmetric Tensor Hierarchies from Superspace Cohomology. 2016. [[arXiv:1607.01402](#)].
- [17] Stephen Randall. A Note on the NATH in Simply Curved Superspace. 2016. (Unpublished).
- [18] Marcus T. Grisaru and W. Siegel. Supergraphity. Part 1. Background Field Formalism. *Nucl.Phys.*, B187:149, 1981. [[INSPIRE entry](#)].
- [19] Chris Beasley and Edward Witten. A Note on fluxes and superpotentials in M theory compactifications on manifolds of G_2 holonomy. *JHEP*, 07:046, 2002. [[arXiv:hep-th/0203061](#)].
- [20] Robbert Dijkgraaf, Sergei Gukov, Andrew Neitzke, and Cumrun Vafa. Topological M-theory as unification of form theories of gravity. *Adv. Theor. Math. Phys.*, 9(4):603–665, 2005. [[arXiv:hep-th/0411073](#)].
- [21] Shuntaro Aoki, Tetsutaro Higaki, Yusuke Yamada, and Ryo Yokokura. Abelian tensor hierarchy in 4D $\mathcal{N} = 1$ conformal supergravity. *JHEP*, 09:148, 2016. [[arXiv:1606.04448](#)].
- [22] Ryo Yokokura. Abelian tensor hierarchy and Chern-Simons actions in 4D $\mathcal{N}=1$ conformal supergravity. 2016. [[arXiv:1609.01111](#)].
- [23] J. Wess and B. Zumino. Superspace Formulation of Supergravity. *Phys. Lett.*, B66:361–364, 1977. [[INSPIRE entry](#)].
- [24] J. Wess and B. Zumino. Superfield Lagrangian for Supergravity. *Phys. Lett.*, B74:51–53, 1978. [[INSPIRE entry](#)].
- [25] Sergei Gukov. Solitons, superpotentials and calibrations. *Nucl. Phys.*, B574:169–188, 2000. [[arXiv:hep-th/9911011](#)].
- [26] Sergei Gukov, Cumrun Vafa, and Edward Witten. CFT’s from Calabi-Yau four folds. *Nucl. Phys.*, B584:69–108, 2000. [Erratum: *Nucl. Phys.*B608,477(2001)] [[arXiv:hep-th/9906070](#)].
- [27] Bobby Samir Acharya and Bill J. Spence. Flux, supersymmetry and M theory on seven manifolds. 2000. [[arXiv:hep-th/0007213](#)].
- [28] Thomas House and Andrei Micu. M-Theory compactifications on manifolds with G_2 structure. *Class. Quant. Grav.*, 22:1709–1738, 2005. [[arXiv:hep-th/0412006](#)].
- [29] Don N. Page. Classical Stability of Round and Squashed Seven Spheres in Eleven-dimensional Supergravity. *Phys. Rev.*, D28:2976, 1983. [[INSPIRE entry](#)].
- [30] Peter G. O. Freund and Mark A. Rubin. Dynamics of Dimensional Reduction. *Phys. Lett.*, B97:233–235, 1980. [[INSPIRE entry](#)].
- [31] D.D. Joyce. *Compact Manifolds with Special Holonomy*. Oxford mathematical monographs. Oxford University Press, 2000.
- [32] Spiro Karigiannis. Deformations of G_2 and $Spin(7)$ structures on manifolds. [[arXiv:math/0301218](#)].